# Magnetic moment of the $X_{1}(2900)$ state in the diquark-antidiquark picture 

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#### Abstract

Motivated by the discovery of fully open-flavor tetraquark states $X_{0}(2900)$ and $X_{1}(2900)$ by the LHCb Collaboration, the magnetic dipole moment of the $X_{1}(2900)$ state with the quantum numbers $J^{P}=1^{-}$is determined in the diquark-antidiquark picture using the lightcone sum rules. The numerical result is obtained as $\mu_{X_{1}}=$ $0.79_{-0.39}^{+0.36} \mu_{N}$. The magnetic moments of hadrons encompasses useful knowledge on the distributions of charge and magnetization their inside, which can be used to better understand their geometrical shapes and quark-gluon organizations. The observation of the $X_{0}(2900)$ and $X_{1}(2900)$ as the first two fully open-flavor multiquark states has opened a new window for investigation of the exotic states. The obtained results in the present study may shed light on the future experimental and theoretical searches on the properties of fully open-flavor multiquark states.


## 1 Introduction

In 2021, the LHCb Collaboration observed two clear peaks in the $D^{-} K^{+}$invariant mass spectrum of the $B^{+} \rightarrow$ $D^{+} D^{-} K^{+}$decay [1,2]. Their spectroscopic parameters are measured to be

$$
\begin{aligned}
X_{0}(2900): M & =2866 \pm 7 \pm 2 \mathrm{MeV} \\
\Gamma & =57 \pm 12 \pm 4 \mathrm{MeV} \\
X_{1}(2900): M & =2904 \pm 5 \pm 1 \mathrm{MeV} \\
\Gamma & =110 \pm 11 \pm 4 \mathrm{MeV}
\end{aligned}
$$

The $D^{-} K^{+}$configuration suggests that the quark constituents of $X_{0}(2900)$ and $X_{1}(2900)$ should be $[\bar{c} \bar{s}][u d]$,

[^0]which means that they are fully open-flavor exotic hadrons. The possible spin-parity quantum numbers of $X_{0}(2900)$ and $X_{1}(2900)$ are estimated to be $J^{P}=0^{+}$and $1^{-}$, respectively.

This observation triggered interesting phenomenological studies on these new resonances in the context of various approaches and models aiming to elucidate their nature, quantum numbers and substructure. These studies are assigned different schemes for these particles: Molecular forms of $\bar{D}^{*} K^{*}$ and $\bar{D}_{1} K$ interactions in Refs. [3-8], the diquark-antidiquark picture in Refs. [3,9-14], and kinematic effects from the triangle singularities in Refs. [15,16]. The production and decay properties of these states were investigated in Refs. [17-20], as well. One can also consult Refs. [21-32] for other pertinent studies on parameters of these states.

Despite the above experimental and theoretical investigations, their properties remain dubious and determination of their exact nature and substructure is still problematic. Indeed, the properties of the $X_{0}(2900)$ and $X_{1}(2900)$ tetraquark states have been suggested differently in different studies. To resolve these ambiguities, their parameters are needed to be further investigated both in theory and experiment. These studies shall include complementary investigations of their spectroscopy as well as their various reactions with other known particles and/or their strong, electromagnetic and weak decay modes.

Inspired by this, in the present study, we are going to consider the interaction of $X_{1}$ (2900) ( $X_{1}$ for short) tetraquark state with photon and examine the magnetic dipole moment of this state within the framework of the light-cone sum rules (LCSR). In the LCSR method, a two-point correlation function is calculated in two different steps. In the first step, it is obtained in terms of hadronic parameters such as form factors, electromagnetic multipole moments, etc., which is
called hadronic representation. In the second step, it is calculated in terms of quark-gluon degrees of freedom, which is called the QCD representation. Then, the correlation functions obtained from these two different ways are related to each other using quark-hadron duality assumption. Finally, the Borel transform and continuum subtraction are performed to suppress the contributions of the possible higher states and continuum. By this way, one obtains the sum rules for the desired physical quantities. In the calculations we use the distribution amplitudes (DAs) of the on-shell photon. When calculating the magnetic dipole moment of $X_{1}$, we will consider that this state has a diquark-antidiquark configuration. There are few studies in the literature where the magnetic dipole moments of the open-flavor exotic states have been investigated, see for instance the Refs. [33-35].

The paper is organized as follows. In Sect. 2, some details of the calculations of the magnetic dipole moment of $X_{1}$ in LCSR method is given. In Sect. 3, we present our numerical results and discussions. In Sect. 4, we discuss obtained results and conclude with brief notes.

## 2 Formalism

As we have mentioned above, at the beginning of the analytic calculations of the magnetic dipole moment, it is necessary to select a sufficient two-point correlation function in the background electromagnetic field, which plays an important role in the LCSR method. It is written as
$\Pi_{\mu \nu}(p, q)=i \int d^{4} x e^{i p \cdot x}\langle 0| \mathcal{T}\left\{J_{\mu}(x) J_{\nu}^{\dagger}(0)\right\}|0\rangle_{\gamma}$,
where $J_{\mu}(x)$ and $\gamma$ represent the interpolating current of $X_{1}$ state and the external electromagnetic field, respectively. We need explicit expression of $J_{\mu}(x)$ to proceed in the calculations. In the diquark-antidiquark picture, $J_{\mu}(x)$ can be written in the following form

$$
\begin{align*}
J_{\mu}^{X_{1}}(x)= & \varepsilon^{a b c} \varepsilon^{a m n}\left[u^{T b}(x) C \gamma_{5} d^{c}(x)\right] \\
& \times\left[\bar{c}^{m}(x) \gamma_{\mu} \gamma_{5} C \bar{s}^{T n}(x)\right], \tag{2}
\end{align*}
$$

where $C$ is the charge conjugation matrix; and $a, b, c, m, n$ are color indices.

In the hadronic representation, two complete sets of the initial and final hadronic states are inserted into the correlation function. By isolating the lowest $X_{1}$ state contribution, we obtain,

$$
\begin{aligned}
\Pi_{\mu \nu}^{H a d}(p, q) & =\frac{\langle 0| J_{\mu}(x)\left|X_{1}\left(p, \varepsilon^{\theta}\right)\right\rangle}{p^{2}-m_{X_{1}}^{2}} \\
& \times\left\langle X_{1}\left(p, \varepsilon^{\theta}\right) \mid X_{1}\left(p+q, \varepsilon^{\delta}\right)\right\rangle_{\gamma}
\end{aligned}
$$

$$
\begin{equation*}
\times \frac{\left\langle X_{1}\left(p+q, \varepsilon^{\delta}\right)\right| J_{v}^{\dagger}(0)|0\rangle}{(p+q)^{2}-m_{X_{1}}^{2}}+\cdots \tag{3}
\end{equation*}
$$

where dots denote the effects of the higher states and continuum. The matrix elements in Eq. (3) are expressed as

$$
\begin{align*}
& \left\langle X_{1}\left(p+q, \varepsilon^{\delta}\right)\right| J_{v}^{\dagger}(0)|0\rangle=\lambda_{X_{1}} \varepsilon_{v}^{\delta}  \tag{4}\\
& \langle 0| J_{\mu}(x)\left|X_{1}\left(p, \varepsilon^{\theta}\right)\right\rangle=\lambda_{X_{1}} \varepsilon_{\mu}^{\theta}  \tag{5}\\
& \left\langle X_{1}\left(p, \varepsilon^{\theta}\right) \mid X_{1}\left(p+q, \varepsilon^{\delta}\right)\right\rangle_{\gamma}=-\varepsilon^{\tau}\left(\varepsilon^{\theta}\right)^{\alpha}\left(\varepsilon^{\delta}\right)^{\beta}\left\{G_{1}\left(Q^{2}\right)\right. \\
& \left.\times(2 p+q)_{\tau} g_{\alpha \beta}+G_{2}\left(Q^{2}\right) g_{\tau \beta} q_{\alpha}-g_{\tau \alpha} q_{\beta}\right) \\
& \left.-\frac{1}{2 m_{X_{1}}^{2}} G_{3}\left(Q^{2}\right)(2 p+q)_{\tau} q_{\alpha} q_{\beta}\right\} \tag{6}
\end{align*}
$$

where $\lambda_{X_{1}}$ is the residue of $X_{1}$; and $\varepsilon_{\mu}^{\theta}, \varepsilon_{v}^{\delta}$ and $\varepsilon^{\tau}$ are the initial and final polarization vectors of the $X_{1}$ and photon polarization vector, respectively. Here, $G_{1}\left(Q^{2}\right), G_{2}\left(Q^{2}\right)$ and $G_{3}\left(Q^{2}\right)$ are electromagnetic form factors, with $Q^{2}=-q^{2}$.

Using Eqs. (3)-(6) and after performing some necessary calculations, the final form of the correlation function is obtained as

$$
\begin{align*}
& \Pi_{\mu \nu}^{H a d}(p, q)=\frac{\varepsilon_{\rho} \lambda_{X_{1}}^{2}}{\left[m_{X_{1}}^{2}-(p+q)^{2}\right]\left[m_{X_{1}}^{2}-p^{2}\right]}\left\{G_{2}\left(Q^{2}\right)\right. \\
& \quad \times\left(q_{\mu} g_{\rho \nu}-q_{\nu} g_{\rho \mu}-\frac{p_{\nu}}{m_{X_{1}}^{2}}\left(q_{\mu} p_{\rho}-\frac{1}{2} Q^{2} g_{\mu \rho}\right)\right. \\
& \left.\quad+\frac{(p+q)_{\mu}}{m_{X_{1}}^{2}}\left(q_{\nu}(p+q)_{\rho}+\frac{1}{2} Q^{2} g_{\nu \rho}\right)-\frac{(p+q)_{\mu} p_{\nu} p_{\rho}}{m_{X_{1}}^{4}} Q^{2}\right) \\
& \quad+\text { other independent structures }\}+\cdots . \tag{7}
\end{align*}
$$

To calculate the magnetic dipole moment, we need to calculate only the form factor $G_{2}\left(Q^{2}\right)$, which is called the magnetic form factor,
$F_{M}\left(Q^{2}\right)=G_{2}\left(Q^{2}\right)$.
At static limit, $Q^{2}=0, F_{M}\left(Q^{2}=0\right)$ is proportional to the magnetic dipole moment $\mu_{X_{1}}$ for real photon:
$\mu_{X_{1}}=\frac{e}{2 m_{X_{1}}} F_{M}\left(Q^{2}=0\right)$.
The correlation function, on the other hand, is determined in terms of the QCD degrees of freedom and the photon distribution amplitudes in the second window called the QCD side. In the QCD representation, we use the Wick's theorem to contract the corresponding quark fields to get the correlation function in terms of the quark propagators and DAs of the photon. After replacing the explicit expression of the interpolating current in the correlation function and applying the Wick's theorem, we get

$$
\begin{align*}
& \Pi_{\mu \nu}^{\mathrm{QCD}}(p, q)=i \int d^{4} x e^{i p x} \varepsilon^{a b c} \varepsilon^{a m n} \varepsilon^{a^{\prime} b^{\prime} c^{\prime}} \varepsilon^{a^{\prime} m^{\prime} n^{\prime}} \\
& \quad \times \operatorname{Tr}\left[\gamma_{5} \widetilde{S}_{u}^{b b^{\prime}}(x) \gamma_{5} S_{d}^{c c^{\prime}}(x)\right] \operatorname{Tr}\left[\gamma_{\mu} \gamma_{5} \widetilde{S}_{s}^{n^{\prime} n}(-x)\right. \\
& \left.\quad \times \gamma_{5} \gamma_{\nu} S_{c}^{m^{\prime} m}(-x)\right]|0\rangle_{\gamma} \tag{10}
\end{align*}
$$

where
$\widetilde{S}_{c(q)}^{i j}(x)=C S_{c(q)}^{i j T}(x) C$,
with $S_{q(c)}(x)$ being the quark propagators. In the $x$-space for the light-quark propagator we use

$$
\begin{align*}
S_{q}(x) & =S_{q}^{\text {free }}-\frac{\langle\bar{q} q\rangle}{12}\left(1-i \frac{m_{q} \not x}{4}\right)-\frac{\langle\bar{q} q\rangle}{192} m_{0}^{2} x^{2}(1 \\
& \left.-i \frac{m_{q} \not x}{6}\right)-\frac{i g_{s}}{32 \pi^{2} x^{2}} G^{\mu \nu}(x)\left[\kappa \sigma_{\mu \nu}+\sigma_{\mu \nu} \not x\right], \tag{11}
\end{align*}
$$

where,
$S_{q}^{\text {free }}=i \frac{\not x}{2 \pi^{2} x^{4}}-\frac{m_{q}}{4 \pi^{2} x^{2}}$.
The charm-quark propagator is given, in terms of the second kind Bessel functions $K_{i}(x)$, as

$$
\begin{align*}
S_{c}(x) & =S_{c}^{\text {free }}-\frac{g_{s} m_{c}}{16 \pi^{2}} \int_{0}^{1} d v G^{\mu \nu}(v x)\left[\left(\sigma_{\mu \nu} \not x+\not \not \sigma_{\mu \nu}\right)\right. \\
& \left.\times \frac{K_{1}\left(m_{c} \sqrt{-x^{2}}\right)}{\sqrt{-x^{2}}}+2 \sigma_{\mu \nu} K_{0}\left(m_{c} \sqrt{-x^{2}}\right)\right] \tag{13}
\end{align*}
$$

where,

$$
\begin{equation*}
S_{c}^{\text {free }}=\frac{m_{c}^{2}}{4 \pi^{2}}\left[\frac{K_{1}\left(m_{c} \sqrt{-x^{2}}\right)}{\sqrt{-x^{2}}}+i \frac{\not x K_{2}\left(m_{c} \sqrt{-x^{2}}\right)}{\left(\sqrt{-x^{2}}\right)^{2}}\right] . \tag{14}
\end{equation*}
$$

The correlation function in QCD representation includes two different contributions: Perturbative and non-perturbative. Practically, the perturbative contribution, in which the photon interacts with one of the quarks perturbatively, can be computed by the replacing one of the light or c-quark propagators by
$S^{\text {free }} \rightarrow \int d^{4} z S^{\text {free }}(x-z) A(z) S^{\text {free }}(z)$,
and the other three propagators with their perturbative or free parts. In the last equation we also use the perturbative parts of the propagators in right hand side as is seen.

For the non-perturbative contribution, in which the photon is radiated at long distances, the correlation function can be
computed by replacing one of the light quark propagators by
$S_{\alpha \beta}^{a b} \rightarrow-\frac{1}{4}\left(\bar{q}^{a} \Gamma_{i} q^{b}\right)\left(\Gamma_{i}\right)_{\alpha \beta}$,
where $\Gamma_{i}=I, \gamma_{5}, \gamma_{\mu}, i \gamma_{5} \gamma_{\mu}, \sigma_{\mu \nu} / 2$, and the remaining light and heavy propagators with their full expressions. We perform all the possible permutations in the perturbative and non-perturbative parts of the correlation function. When Eq. (16) is employed in computation of the nonperturbative effects, we observe that matrix elements of the forms $\langle\gamma(q)| \bar{q}(x) \Gamma_{i} q(0)|0\rangle$ and $\langle\gamma(q)| \bar{q}(x) \Gamma_{i} G_{\mu \nu} q(0)|0\rangle$ appear. These matrix elements are written in terms of the photon distribution amplitudes (see Ref. [36]). Using these matrix elements in terms of photon's DAs and the expressions of the propagators given above, the QCD representation of the correlation function in coordinate space is obtained. We perform Fourier transformation to carry the calculations to the momentum space.

The LCSR for the magnetic dipole moment of $X_{1}$ state can be acquired by matching the functions $\Pi_{\mu \nu}^{Q C D}(p)$ and $\Pi_{\mu \nu}^{H a d}(p)$ from both the QCD and hadronic sides. We choose the $\varepsilon_{\nu} q_{\mu}$ structure and equate the coefficients of this structure from both sides to each other. We apply a double Borel transformation with respect to $-p^{2}$ and $-(p+q)^{2}$ and also continuum subtraction procedure based on the standard prescriptions of the method in order to suppress the contributions of the higher states and continuum (fore details see for instance Refs. [37,38]). As a results, we get the desired LCSR for the magnetic dipole moment as
$\mu_{X_{1}}=\frac{e^{\frac{m_{X_{1}}^{2}}{M^{2}}}}{\lambda_{X_{1}}^{2}} \Delta\left(M^{2}, s_{0}\right)$,
where the explicit expression of $\Delta\left(M^{2}, s_{0}\right)$ function, which represents the final form of the QCD side of the calculations, is presented in the Appendix. In obtaining the last result, we use $\frac{1}{M^{2}}=\frac{1}{M_{1}^{2}}+\frac{1}{M_{2}^{2}}$ with $M_{1}^{2}$ and $M_{2}^{2}$ being the Borel parameters in the initial and final channels, respectively. We set $M_{1}^{2}=M_{2}^{2}=2 M^{2}$ as the initial and final particles are the same. We will fix $M^{2}$ and continuum threshold ( $s_{0}$ ), coming from the continuum subtraction procedure and appearing inside the $\Delta\left(M^{2}, s_{0}\right)$ function, based on the standard criteria of the method in next section.

## 3 Numerical analysis

In this section, we present our numerical prediction for the magnetic dipole moment of the $X_{1}$ state. To get our numerical results, the main input parameters are the photon distribution amplitudes. The distribution amplitudes of the pho-
ton, which depend on various non-perturbative parameters are borrowed from Ref. [36]. We use the values of some other input parameters as follow: $m_{u}=m_{d}=0, m_{s}=$ $96_{-4}^{+8} \mathrm{MeV}, m_{c}=(1.275 \pm 0.025) \mathrm{GeV}, m_{X_{1}}=2904 \pm$ $5 \pm 1 \mathrm{MeV}, f_{3 \gamma}=-0.0039 G e V^{2}[36],\langle\bar{s} s\rangle=0.8\langle\bar{u} u\rangle$ with $\langle\bar{u} u\rangle=(-0.24 \pm 0.01)^{3} \mathrm{GeV}^{3}$ [39], $m_{0}^{2}=0.8 \pm 0.1 \mathrm{GeV}^{2}$ [39], $\left\langle\frac{\alpha_{s}}{\pi} G^{2}\right\rangle=(0.012 \pm 0.004) \mathrm{GeV}^{4}[40]$ and $\lambda_{X_{1}}=m_{X_{1}} f_{X_{1}}$ with $f_{X_{1}}=(2.1 \pm 0.4) \times 10^{-3} \mathrm{GeV}^{4}$ [14].

In the LCSR method, in addition to the DAs and above parameters, there are two extra auxiliary parameters as mentioned before: The Borel mass parameter $M^{2}$ and the continuum threshold $s_{0}$. The continuum threshold is not totally arbitrary but it shows the energy scale at which, the excited states and continuum begin to contribute to the correlation function. The physical quantities like the magnetic dipole moment are expected to be independent of these helping parameters. In practice, however, there appear some residual dependence on these parameters which are entered as the uncertainties to the presented results. To find the working intervals of these auxiliary parameters, we demand that both the continuum and higher states contributions have to be sufficiently suppressed and series of the operator product expansion (OPE) in QCD side converge. To this end, in technique language, we define the pole contribution (PC) as
$\mathrm{PC}=\frac{\Delta\left(M^{2}, s_{0}\right)}{\Delta\left(M^{2}, \infty\right)}$,
and require that it should exceed $20 \%$ of the total contributions, which is typical for the multiquark states. We also demand that the series of light-cone expansion converges and contributions of the higher twist and higher condensate terms are less than $10 \%$ of the total contribution. These considerations lead to the following working windows for $M^{2}$ and $s_{0}$ :

$$
\begin{aligned}
& 3.0 \mathrm{GeV}^{2} \leq M^{2} \leq 3.5 \mathrm{GeV}^{2} \\
& 11.0 \mathrm{GeV}^{2} \leq s_{0} \leq 12.5 \mathrm{GeV}^{2}
\end{aligned}
$$

Our numerical analyses show that, by considering these working regions for the auxiliary parameters, for the magnetic dipole moment of the $X_{1}$ state $P C$ varies within the interval $30 \% \leq P C \leq 64 \%$ corresponding to the upper and lower limits of the Borel mass parameter. When we analyze the OPE convergence, we see that the contribution of the higher twist and higher dimensional terms in OPE is $2 \%$ of the total and the series show a good convergence. It is worth mentioning that the above interval for the continuum threshold corresponds to $s_{0} \simeq\left(m_{X_{1}}+0.5_{-0.1}^{+0.1}\right)^{2} \mathrm{GeV}^{2}$, which is typical in hadronic spectrum. Therefore, the chosen working windows for $M^{2}$ and $s_{0}$ well fulfill the requirements of the light-cone sum rules method.

In Fig. 1, we depict the dependence of the magnetic dipole moment of the $X_{1}$ state on $M^{2}$ at three fixed values of $s_{0}$. As


Fig. 1 The magnetic dipole moment of the $X_{1}$ state versus $M^{2}$ at three fixed values of $s_{0}$
one can see from this figure, the magnetic dipole moment show a good stability with respect to the variation of the Borel mass parameter. Although the dependence on $s_{0}$ is considerable, however, it remains within the limits allowed by the method to calculate the magnetic dipole moment.

Considering all the input parameters, DAs of the photon and the working intervals of auxiliary parameters, our prediction for the magnetic dipole moment of the $X_{1}$ state, both in its natural unit $\left(\frac{e}{2 m_{X_{1}}}\right)$ and nuclear magneton $\left(\mu_{N}=\frac{e}{2 m_{N}}\right)$, is
$\mu_{X_{1}}=2.43_{-1.21}^{+1.13} \frac{e}{2 m_{X_{1}}}=0.79_{-0.39}^{+0.36} \mu_{N}$.
The order of magnitude for the magnetic dipole moment shows that measurement of $\mu_{X_{1}}$ is accessible in the future experiments.

## 4 Summary and conclusion

Motivated by the discovery of fully open-flavor tetraquark states $X_{0}(2900)$ and $X_{1}(2900)$ by the LHCb Collaboration, the magnetic dipole moment of the $X_{1}(2900)$ tetraquark state have been determined using the light-cone sum rules assigning the diquark-antidiquark structure with the quantum numbers $J^{P}=1^{-}$for this state. The magnetic moments of hadrons encompasses useful knowledge about the distribution of charge and magnetization their inside, which helps us to understand their nature, quark-gluon organization and geometrical shape. The existing theoretical predictions on the spectroscopic parameters of $X_{1}(2900)$ tetraquark and their comparison with the experimental data have also given rise to different assignments on the internal structure of this state. Calculations of electromagnetic parameters of the exotic states like their magnetic dipole moment can be useful in establishing the nature of these states. The observation of
the $X_{0}(2900)$ and $X_{1}(2900)$ by LHCb as the first two fully open-flavor multiquark states has opened a new platform for investigation of the exotic states. More experimental and theoretical research is required to fully understand the properties of this class of hadrons. The magnitude obtained for the magnetic dipole moment of $X_{1}(2900)$ shows a possibility to measure it in future experiments. The obtained result in the present study may be useful for analyses of the future experimental data on parameters of fully open-flavor multiquark states.

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Data Availability Statement This manuscript has no associated data or the data will not be deposited. [Authors' comment: All the numerical and mathematical data have been included in the paper and we have no other data regarding this paper.]

## Appendix: explicit expression of $\Delta\left(M^{2}, s_{0}\right)$

In this appendix, we present the explicit expression of the function $\Delta\left(M^{2}, s_{0}\right)$ entering into the sum rule for the magnetic moment of the $X_{1}$ state. It is obtained as

$$
\begin{aligned}
& \Delta\left(M^{2}, s_{0}\right)=-\frac{e_{c}}{36864 m_{c}^{2} \pi^{6}}\left[3 m_{c}^{12}(3 I[-6,4]-4 I[-5,3])\right. \\
& -48 m_{c}^{10}(I[-5,4]+I[-4,3])-48 m_{c}^{9} m_{s} I[-4,3] \\
& +3 m_{c}^{8}\left(30 I[-4,4]+P_{1}(I[-4,2]-2 I[-3,1])\right. \\
& \left.+96 m_{s} P_{3} \pi^{2}(I[-4,2]-2 I[-3,1])-24 I[-3,3]\right) \\
& +144 m_{c}^{7}\left(4 P_{3} \pi^{2} I[-3,2]-m_{s} I[-3,3]\right) \\
& -12 m_{c}^{6}\left(-32 m_{0}^{2} m_{s} P_{3} \pi^{2} I[-3,1]\right. \\
& +512 P_{2}^{2} \pi^{4} I[-3,1]+P_{1} I[-3,2] \\
& +96 m_{s} P_{3} \pi^{2} I[-3,2]+6 I[-3,4]+P_{1} I[-2,1] \\
& \left.+96 m_{s} P_{3} \pi^{2} I[-2,1]+4 I[-2,3]\right) \\
& -12 m_{c}^{5}\left(48 P_{3} \pi^{2}\left(m_{0}^{2} I[-2,1]+2 I[-2,2]\right)\right. \\
& \left.+m_{s}\left(P_{1} I[-2,1]+12 I[-2,3]\right)\right) \\
& +3 m_{c}^{4}\left(64 m_{0}^{2} m_{s} P_{3} \pi^{2} I[-2,1]\right. \\
& -1024 P_{2}^{2} \pi^{4} I[-2,1]+3 P_{1} I[-2,2]+288 m_{s} P_{3} \pi^{2} I[-2,2] \\
& +7 I[-2,4]-2 P_{1} I[-1,1]-192 m_{s} P_{3} \pi^{2} I[-1,1] \\
& -4 I[-1,3])+8\left(-192 P_{2}^{2} \pi^{4}\left(m_{0}^{2} I[0,0]+2 I[0,1]\right)\right. \\
& \left.+m_{s} P_{3} \pi^{2}\left(P_{1} I[0,0]+24 m_{0}^{2} I[0,1]\right)\right)-192 m_{c}^{2} I[0,3] \\
& -16 m_{c}\left(P_{1} P_{3} \pi^{2} I[0,0]+24 m_{s}\left(-8 P_{2}^{2} \pi^{4} I[0,0]+I[0,3]\right)\right) \\
& +12 m_{c}^{3}\left(48 P_{3} \pi^{2}(I[-1,2]\right. \\
& \left.\left.\left.-m_{0}^{2} I[1,0]\right)-m_{s}\left(4 I[-1,3]+P_{1} I[1,0]\right)\right)\right]
\end{aligned}
$$

$$
\begin{aligned}
& +\frac{e_{d}}{442368 m_{c}^{2} \pi^{6}}\left[f _ { 3 \gamma } \pi ^ { 2 } \left(-11 P_{1}\left(m_{c}^{6} I[-3,1]\right.\right.\right. \\
& \left.+3 m_{c}^{4} I[-2,1]+2 I[0,1]\right) I_{1}[\mathcal{A}]+12\left(5 m_{c}^{10} I[-5,3]\right. \\
& -12 m_{c}^{8} I[-4,3]-36 m_{c}^{7} m_{s} I[-3,2] \\
& +m_{c}^{6}\left(48 m_{s} P_{3} \pi^{2} I[-3,1]-9 I[-3,3]\right) \\
& +24 m_{c}^{5}\left(8 P_{3} \pi^{2} I[-2,1]\right. \\
& \left.+3 m_{s} I[-2,2]\right)-2 m_{c}^{4}\left(24 m_{s} P_{3} \pi^{2} I[-2,1]\right. \\
& +5 I[-2,3])-36 m_{c}^{3} m_{s} I[-1,2]+48 m_{0}^{2} m_{c} P_{3} \pi^{2} I[0,0] \\
& +8 m_{s} P_{3} \pi^{2}\left(m_{0}^{2} I[0,0]-12 I[0,1]\right) \\
& \left.-20 I[0,3]) I_{1}[\mathcal{V}]\right)+4\left(1152 m_{c}^{4} P_{2}^{2} \pi^{4} I_{4}[\mathcal{S}] I[-2,1]\right. \\
& -6 m_{c}^{5} P_{1}\left(m _ { c } \left(2 m_{c}\left(m_{c}+m_{s}\right) I[-3,1]\right.\right. \\
& \left.+I[-3,2])+\left(3 m_{c}+2 m_{s}\right) I[-2,1]\right) \\
& +9 m_{c}^{4} P_{1} I[-2,2]-6 m_{c}^{4} P_{1} I[-1,1] \\
& -6 m_{c}^{3} m_{s} P_{1} I[-1,1]+12 m_{c}^{3} m_{s} P_{1} I[0,0] \\
& +16 m_{c} P_{1} P_{3} \pi^{2} I[0,0]+8 m_{s} P_{1} P_{3} \pi I[0,0] \\
& -1152 m_{c} m_{s} P_{2}^{2} \pi^{4} I_{4}[\mathcal{S}] I[0,0]-48 m_{c} m_{s} P_{1} I[0,1] \\
& +1152 P_{2}^{2} \pi^{4} I_{4}[\mathcal{S}] I[0,1]-144 P_{2}^{2} \pi^{4}\left(6 m_{c}^{6} I[-3,1]\right. \\
& \left.+4 m_{c}^{4} I[-2,1]+m_{0}^{2} I[0,0]+2 I[0,1]\right) I_{1}[\mathcal{S}] \\
& +9 P_{1} I[0,2]+18 m_{c}^{3} m_{s} P_{1} I[1,0] \\
& +2 f_{3 \gamma} P_{1} \pi^{2}\left(3 m_{c}^{6} I[-3,1]\right. \\
& \left.+2 m_{c}^{4} I[-2,1]+3 I[0,1]+2 m_{c}^{3} m_{s} I[1,0]\right) I_{6}\left[\psi_{a}\right] \\
& +4 f_{3 \gamma} P_{1} \pi^{2}\left(3 m_{c}^{6} I[-3,1]\right. \\
& \left.\left.\left.+2 m_{c} m_{s} I[0,0]-I[0,1]\right) \psi^{a}\left[u_{0}\right]\right)\right] \\
& +\frac{e_{u}}{442368 m_{c}^{2} \pi^{6}}\left[f _ { 3 \gamma } \pi ^ { 2 } \left(1 2 \left(5 m_{c} 10 I[-5,3]\right.\right.\right. \\
& -12 m_{c}^{8} I[-4,3]-36 m_{c}^{7} m_{s} I[-3,2] \\
& +m_{c}^{6}\left(48 m_{s} P_{3} \pi^{2} I[-3,1]\right. \\
& -9 I[-3,3])+24 m_{c}^{5}\left(8 P_{3} \pi^{2} I[-2,1]+3 m_{s} I[-2,2]\right) \\
& -2 m_{c}^{4}\left(24 m_{s} P_{3} \pi^{2} I[-2,1]+5 I[-2,3]\right) \\
& -36 m_{c}^{3} m_{s} I[-1,2]+48 m_{0}^{2} m_{c} P_{3} \pi^{2} I[0,0] \\
& \left.+8 m_{s} P_{3} \pi^{2}\left(m_{0}^{2} I[0,0]-12 I[0,1]\right)-44 I[0,3]\right) I_{2}[\mathcal{V}] \\
& +P_{1}\left(1 1 \left(m_{c}^{6} I[-3,1]+3 m_{c}^{4} I[-2,1]\right.\right. \\
& +2 I[0,1]) I_{2}[\mathcal{A}]+8\left(3 m_{c}^{6} I[-3,1]\right. \\
& +2 m_{c}^{4} I[-2,1]+3 I[0,1] \\
& \left.+2 m_{c}^{3} m_{s} I[1,0]\right) I_{6}\left[\psi^{a}\right] \\
& +16\left(3 m_{c}^{6} I[-3,1]+2 m_{c} m_{s} I[0,0]\right. \\
& \left.\left.-I[0,1]) \psi^{a}\left[u_{0}\right]\right)\right)-24 m_{c}^{5}\left(3 m_{c}+2 m_{s}\right) P_{1} I[-2,1] \\
& -4608 P_{2}^{2} \pi^{4}\left(m_{c}^{4} I[-2,1]-m_{c} m_{s} I[0,0]+I[0,1]\right) I_{3}[\mathcal{S}] \\
& +576 P_{2}^{2} \pi^{4}\left(6 m_{c}^{6} I[-3,1]+4 m_{c}^{4} I[-2,1]+m_{0}^{2} I[0,0]\right. \\
& +2 I[0,1]) I_{2}[\mathcal{S}]+4 P_{1}\left(m_{c}^{4}(9 I[-2,2]-6 I[-1,1])\right. \\
& +8 m_{s} P_{3} \pi^{2} I[0,0]+16 m_{c}\left(P_{3} \pi^{2} I[0,0]-3 m_{s} I[0,1]\right)
\end{aligned}
$$

$\left.\left.+9 I[0,2]-6 m_{c}^{3} m_{s}(I[-1,1]-2 I[0,0]-3 I[1,0])\right)\right]$
$-\frac{e_{s}}{221184 m_{c}^{2} \pi^{4}}\left[m_{c}^{12}(6 I[-6,4]+8 I[-5,3])\right.$
$-24 m_{c} 10(I[-5,4]-I[-4,3])$
$+m_{c}^{8}(P 1 I[-4,2]+30 I[-4,4]$
$\left.+2 P_{1} I[-3,1]+24 I[-3,3]\right)$
$+2 m_{c}^{6}\left(512 P_{2}^{2} \pi^{4} I[-3,1]-P_{1} I[-3,2]\right.$
$\left.-6 I[-3,4]+P_{1} I[-2,1]+4 I[-2,3]\right)$
$+1024 P_{2}^{2} \pi^{4} I[0,1]-3 P_{1} I[0,2]+64 m_{c}^{2} I[0,3]$
$+P_{3}\left(864\left(m_{c}^{6} I[-3,2]-m_{c}^{4} I[-2,2]\right) \mathcal{A}\left[u_{0}\right]\right.$
$-432 m_{c}^{2}\left(I_{4}[\mathcal{S}]+I_{4}\left[\mathcal{T}_{1}\right]+I_{4}\left[\mathcal{T}_{2}\right]-I_{4}\left[\mathcal{T}_{3}\right]\right.$
$\left.-I_{4}\left[\mathcal{T}_{4}\right]-I_{4}[\tilde{\mathcal{S}}]\right)\left(m_{c}^{4} I[-3,2]\right.$
$\left.-2 m_{c}^{2} I[-2,2]+I[-1,2]\right)$
$+P_{1}\left(23 I_{4}[\mathcal{S}]+23 I_{4}\left[\mathcal{T}_{1}\right]+23 I_{4}\left[\mathcal{T}_{2}\right]\right.$
$\left.-12\left(I_{4}\left[\mathcal{T}_{3}\right]+I_{4}\left[\mathcal{T}_{4}\right]+I_{4}[\tilde{\mathcal{S}}]\right)\right) I[0,0]$
$+24\left(36 m_{c}^{4}\left(m_{c}^{4} I[-4,2]-2 m_{c}^{2} I[-3,2]\right.\right.$
$\left.\left.+I[-2,2])-P_{1} I[0,0]\right) I_{6}\left[h_{\gamma}\right]\right)$
$-f_{3 \gamma}\left(144 m_{c}^{8} I_{1}[\mathcal{A}] I[-4,3]\right.$
$+144 m_{c}^{8} I_{1}[\mathcal{V}] I[-4,3]+288 m_{c}^{6} I_{1}[\mathcal{A}] I[-3,3]$
$+288 m_{c}^{6} I_{1}[\mathcal{V}] I[-3,3]$
$+23 m_{c}^{4} P_{1} I_{1}[\mathcal{A}] I[-2,1]+23 m_{c}^{4} P_{1} I_{1}[\mathcal{V}] I[-2,1]$
$+144 m_{c}^{4} I_{1}[\mathcal{A}] I[-2,3]+144 m_{c}^{4} I_{1}[\mathcal{V}] I[-2,3]$
$+23 P_{1} I_{1}[\mathcal{A}] I[0,1]+23 P_{1} I_{1}[\mathcal{V}] I[0,1]$
$+576 I_{1}[\mathcal{A}] I[0,3]+576 I_{1}[\mathcal{V}] I[0,3]$
$+24\left(m_{c}^{6}\left(12 m_{c}^{4} I[-5,3]+24 m_{c}^{2} I[-4,3]\right.\right.$
$\left.+P_{1} I[-3,1]+12 I[-3,3]\right)-P_{1} I[0,1]$
$+48 I[0,3]) I_{6}\left[\psi^{a}\right]$
$+96\left(m_{c}^{6}\left(12 m_{c}^{4} I[-5,3]+24 m_{c}^{2} I[-4,3]\right.\right.$
$\left.+P_{1} I[-3,1]+12 I[-3,3]\right)+P_{1} I[0,1]$
$+48 I[0,3]) I_{6}\left[\varphi_{\gamma}\right]$
$+576 m_{c}^{10} I[-5,3] \psi^{a}\left[u_{0}\right]+48 m_{c}^{6} P_{1} I[-3,1] \psi^{a}\left[u_{0}\right]$
$-576 m_{c}^{6} I[-3,3] \psi^{a}\left[u_{0}\right]+48 P_{1} I[0,1] \psi^{a}\left[u_{0}\right]$
$+48\left(m_{c}^{6}\left(12 m_{c}^{4} I[-5,3]\right.\right.$
$+24 m_{c}^{2} I[-4,3]+P_{1} I[-3,1]$
$+12 I[-3,3])+P_{1} I[0,1]$
$\left.\left.+48 I[0,3]) \varphi_{\gamma}\left[u_{0}\right]\right)\right]$,
where $P_{1}=\left\langle g_{s}^{2} G^{2}\right\rangle$ is gluon condensate, $P_{2}=\langle\bar{q} q\rangle$ stands for $\mathrm{u} / \mathrm{d}$ quark condensate, and $\left.P_{3}=\langle\bar{s}\rangle\right\rangle$ represents the
s-quark condensate. The functions $I[n, m], I_{1}[\mathcal{A}], I_{2}[\mathcal{A}]$, $I_{3}[\mathcal{A}], I_{4}[\mathcal{A}], I_{5}[\mathcal{A}]$ and $I_{6}[\mathcal{A}]$ are defined as

$$
\begin{aligned}
I[n, m] & =\int_{m_{c}^{2}}^{s_{0}} d s \int_{m_{c}^{2}}^{s} d l e^{-s / M^{2}} \frac{(s-l)^{m}}{l^{n}} \\
I_{1}[\mathcal{A}] & =\int D_{\alpha_{i}} \int_{0}^{1} d v \mathcal{A}\left(\alpha_{\bar{q}}, \alpha_{q}, \alpha_{g}\right) \delta^{\prime}\left(\alpha_{q}+\bar{v} \alpha_{g}-u_{0}\right), \\
I_{2}[\mathcal{A}] & =\int D_{\alpha_{i}} \int_{0}^{1} d v \mathcal{A}\left(\alpha_{\bar{q}}, \alpha_{q}, \alpha_{g}\right) \delta^{\prime}\left(\alpha_{\bar{q}}+v \alpha_{g}-u_{0}\right), \\
I_{3}[\mathcal{A}] & =\int_{0}^{1} d u A(u) \delta^{\prime}\left(u-u_{0}\right), \\
I_{4}[\mathcal{A}] & =\int D_{\alpha_{i}} \int_{0}^{1} d v \mathcal{A}\left(\alpha_{\bar{q}}, \alpha_{q}, \alpha_{g}\right) \delta\left(\alpha_{q}+\bar{v} \alpha_{g}-u_{0}\right), \\
I_{5}[\mathcal{A}] & =\int D_{\alpha_{i}} \int_{0}^{1} d v \mathcal{A}\left(\alpha_{\bar{q}}, \alpha_{q}, \alpha_{g}\right) \delta\left(\alpha_{\bar{q}}+v \alpha_{g}-u_{0}\right), \\
I_{6}[\mathcal{A}] & =\int_{0}^{1} d u A(u),
\end{aligned}
$$

where $\mathcal{A}$ stands for the corresponding photon DAs and $D_{\alpha_{i}}$ is the measure defined as

$$
\int \mathcal{D} \alpha_{i}=\int_{0}^{1} d \alpha_{\bar{q}} \int_{0}^{1} d \alpha_{q} \int_{0}^{1} d \alpha_{g} \delta\left(1-\alpha_{\bar{q}}-\alpha_{q}-\alpha_{g}\right)
$$

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